

The total cross-sections for the photoeffect for K-shell bound electrons and pair production with the created electron in the ground state for photon energies above 1 MeV

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Considering the contributions of the main term of the relativistic Coulombian Green function given by Hostler to the second order of S matrix element and taking into account only the large components of the Dirac spinor of the ground state we obtain the imaginary part of the Rayleigh amplitudes in terms of elementary functions. Thereby simple and high accurate formulae for the total cross-sections for photoeffect and pair production with the electron created in the K-shell are obtained *via* the optical theorem. Comparing the predictions given by our formulae with the full relativistic numerical calculations of Kissel *et al* [Phys. Rev. A 22, 1970 (1980)] and Scofield [LLRL, Internal Report, 1973], a good agreement is found for photon energies above the pair production threshold up to 5 MeV for any Z elements. We present our numerical results for the total photoeffect and pair production cross-sections, for various photon energies for the K-shell of Ag and Pb.

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I. INTRODUCTION

The S matrix element for Rayleigh scattering of an initial photon with momentum $\vec{k}_1 = \omega \vec{v}_1$ and polarization vector \vec{s}_1 , and a final photon with momentum $\vec{k}_2 = \omega \vec{v}_2$ and polarization vector \vec{s}_2 , by a K-shell bound electron, if we take into account only the large components of the Dirac spinor of the ground state and only the main term of the relativistic coulombian Green function[1], is:[4]

$$\mathcal{M}_R = M(\omega, \theta) (\vec{s}_1 \vec{s}_2) + N(\omega, \theta) (\vec{s}_1 \vec{v}_2) (\vec{s}_2 \vec{v}_1) \quad (1.1)$$

where θ is the photon scattering angle, and the invariant amplitudes M and N are:

$$M(\omega, \theta) = \mathcal{O} - P(\Omega_1, \theta) - P(\Omega_2, \theta) \quad (1.2)$$

$$N(\omega, \theta) = Q(\Omega_1, \theta) + Q(\Omega_2, \theta) \quad (1.3)$$

with the Rayleigh scattering amplitudes given by:

$$P(\Omega, \theta) = \frac{2^7 \lambda^5 X^3 \omega \pm \omega_{pp}}{d^4(\Omega)} \frac{F_1(2 - \tau; 2, 2; 3 - \tau; x_1, x_2)}{2 - \tau} \quad (1.4)$$

$$Q(\Omega, \theta) = \frac{2^{11} \lambda^5 X^3 \omega^2 \pm \omega_{pp}}{d^5(\Omega)} \frac{F_1(3 - \tau; 3, 3; 4 - \tau; x_1, x_2)}{3 - \tau} \quad (1.5)$$

where

$$\Omega_1 = \gamma m + \omega, \Omega_2 = \gamma m - \omega = -|\Omega_2|,$$

$$\gamma = (1 - \alpha^2 Z^2)^{\frac{1}{2}}, \lambda = \alpha Z m, \omega_{pp} = (1 + \gamma) m,$$

$$\tau_j = \frac{\alpha Z \Omega_j}{X_j}, X_j = m^2 - \Omega_j^2 \text{ with } \text{Re}[X_j] > 0, j = 1, 2$$

The function \mathcal{O} is the atomic form factor, while $F_1(a; b_1, b_2; c; x_1, x_2)$ is the Appell hypergeometric function of four parameters and two complex variables given by the relationships:

$$x_1 x_2 = p = \left[\frac{d^*(\Omega_1)}{d(\Omega_1)} \right]^2 = \xi^2 \quad (1.6)$$

$$x_1 + x_2 = s = 2\xi - \frac{16X^2 \omega^2 \sin^2 \frac{\theta}{2}}{d^2(\Omega_1)} = \xi^2 \quad (1.7)$$

$$\begin{aligned} \text{with } d(\Omega_j) &= 2m(\pm\gamma\omega - \alpha^2 Z^2 m + \alpha Z X_j), \\ d^*(\Omega_j) &= 2m(\pm\gamma\omega - \alpha^2 Z^2 m - \alpha Z X_j) \end{aligned} \quad (1.8)$$

In the equations (1.4)-(1.5) and (1.8) the upper sign corresponds to the case $\Omega = \Omega_1$, while the lower sign corresponds to the case $\Omega = \Omega_2$.

Obviously, in the case of the forward scattering, the amplitude $N(\omega, \theta)$ is no longer present in the Rayleigh amplitude. Also, the imaginary part of the amplitude \mathcal{M}_R is given only by the terms $P(\Omega_1, 0)$ and $P(\Omega_2, 0)$ which are present in the expressions of the photoeffect and the pair production cross-sections respectively.

II. THE TOTAL CROSS-SECTION OF THE PHOTOELECTRIC EFFECT AND PAIR PRODUCTION IN THE CASE OF K SHELL

According to the optical theorem, the imaginary part of the Rayleigh amplitude for forward scattering allows to get the total photoeffect cross-section *per* K-shell electron:

$$\sigma_{ph} = \frac{4\pi}{\alpha} \frac{m}{\omega} r_0^2 |\text{Im}[P(\Omega_1, 0)]| = \frac{16\pi^2}{3} r_0^2 m^2 \alpha^5 Z^6 \frac{\omega + \omega_{pp}}{2m} \frac{\Omega_1 |X_1^2|}{\omega^5} (1 + |\tau_1|^2) \frac{(-\xi_1)^{\tau_1}}{e^{\pi|\tau_1|} - e^{-\pi|\tau_1|}} \quad (2.1)$$

In a similar way we get the total pair production cross-section *per* K-shell electron:

$$\sigma_{pp} = \frac{4\pi}{\alpha} \frac{m}{\omega} r_0^2 |\text{Im}[P(\Omega_2, 0)]| = \frac{16\pi^2}{3} r_0^2 m^2 \alpha^5 Z^6 \frac{\omega - \omega_{pp}}{2m} \frac{|\Omega_2| |X_2^2|}{\omega^5} (1 + |\tau_2|^2) \frac{(-\xi_2)^{\tau_2}}{e^{\pi|\tau_2|} - e^{-\pi|\tau_2|}} \quad (2.2)$$

We point out that: $(-\xi_1)^{\tau_1} = e^{-|\tau_1|\chi_1}$, with

$$\chi_1 = \pi - 2 \arctan \left(\frac{\alpha Z |X_1|}{\gamma \omega - \alpha^2 Z^2 m} \right); \omega > \frac{\alpha^2 Z^2 m}{\gamma} \quad (2.3)$$

and $(-\xi_2)^{\tau_2} = e^{-|\tau_2|\chi_2}$, with

$$\chi_2 = \pi - 2 \arctan \left(\frac{\alpha Z |X_2|}{\gamma \omega + \alpha^2 Z^2 m} \right); \omega > \omega_{pp} \quad (2.4)$$

III. NUMERICAL RESULTS AND CONCLUSIONS

Using our analytical formulae for the cross sections for K-shell electrons we get the numerical numerical results in Table I, Table II, and figure 1 for the whole K-shell.

TABLE I: Photoeffect and pair production cross sections for Ag K-shell.

| Energy (keV) | Pair production cross section (mb) | Photoeffect cross section (mb) | Cross sections ratio |
|--------------|------------------------------------|--------------------------------|-----------------------|
| 1000 | 9.145x10 ⁻⁷ | 493.836 | 5.399x10 ⁸ |
| 1100 | 0.0875335 | 401.376 | 4585.39 |
| 1249 | 0.851678 | 306.868 | 360.309 |
| 1500 | 3.11888 | 211.705 | 67.878 |
| 2000 | 7.16782 | 122.626 | 17.107 |
| 2754 | 9.96235 | 70.2819 | 7.054 |
| 3000 | 10.3218 | 61.0716 | 5.916 |
| 3500 | 10.6112 | 47.8107 | 4.505 |
| 4000 | 10.5516 | 38.99 | 3.695 |
| 4807 | 10.1208 | 29.7828 | 2.942 |
| 5000 | 9.98901 | 28.1578 | 2.818 |
| 5500 | 9.62702 | 24.6344 | 2.558 |
| 6000 | 9.25496 | 21.8602 | 2.362 |
| 6500 | 8.88765 | 19.6248 | 2.208 |
| 7000 | 8.53296 | 17.7887 | 2.084 |
| 7500 | 8.1948 | 16.2559 | 1.983 |

As it may be observed from equations (2.3) and (2.4), for higher photon energies the arguments of the exponentials, χ_1 and χ_2 , have the same limit, $\pi - 2 \arctan[\frac{\alpha Z}{\gamma}]$, so that the two cross sections will become closer and closer as the photon energy increases.

Comparing the predictions given by our formulae with the full relativistic numerical calculations of Kissel *et al* [2] and Scofield[3], a good agreement is found for photon energies above the pair production threshold up to 5 MeV for any Z elements.

The good agreement of our calculation with the full relativistic results shows that, for the presented energies

regime, the main relativistic kinematics terms are canceled by some retardation and multipoles terms, and the spin effects are small.

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TABLE II: Photoeffect and pair production cross sections for Pb K-shell.

| Energy (keV) | Pair production cross section (mb) | Photoeffect cross section (mb) | Cross sections ratio |
|--------------|---------------------------------------|-----------------------------------|-------------------------|
| 1000 | 0.11370 | 5099.07 | 44846.5 |
| 1100 | 2.36329 | 4151.89 | 1756.82 |
| 1249 | 12.2336 | 3181.11 | 260.03 |
| 1500 | 37.1725 | 2200.44 | 59.195 |
| 2000 | 79.4739 | 1278.86 | 16.091 |
| 2754 | 107.831 | 735.099 | 6.817 |
| 3000 | 111.338 | 639.19 | 5.741 |
| 3500 | 113.924 | 500.947 | 4.397 |
| 4000 | 112.949 | 408.876 | 3.619 |
| 4807 | 108.012 | 312.65 | 2.894 |
| 5000 | 106.549 | 295.652 | 2.774 |
| 5500 | 102.571 | 258.781 | 2.522 |
| 6000 | 98.517 | 229.731 | 2.331 |
| 6500 | 94.5392 | 206.311 | 2.182 |
| 7000 | 90.7119 | 187.065 | 2.062 |
| 7500 | 87.0731 | 170.993 | 1.963 |

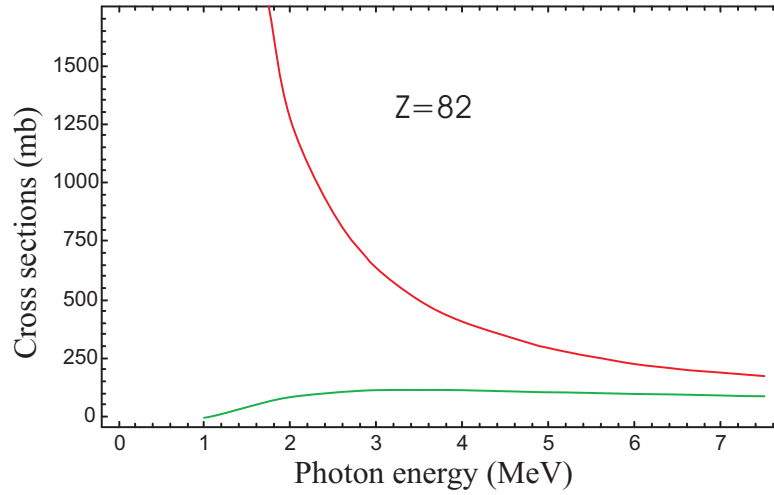


FIG. 1: Photoeffect (red) and pair production (green) cross sections for Pb K-shell

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